X-ray Astronomy: towards the next 50 years!

Milan, October 1-5, 2012 National Museum of Science and Technology Leonardo da Vinci

Scientific Organizing Committee

M. Arnaud (Laboratoire AIM, IRFU/Service dAstrophysique, Saclay, France)

- X. Barcons (Instituto de Fisica de Cantabria, Santander, Spain)
- S. Bowyer (University of California, Berkeley, USA)
- C. Canizares (Massachusetts Institute of Technology, Cambridge, USA)
- R. Giacconi (The Johns Hopkins University, Baltimore, USA)
- F. Harrison (California Institute of Technology, Pasadena, USA)
- S. Holt (Franklin W. Olin College of Engineering, Needham, USA)
- T. Maccacaro (INAF-Osservatorio Astronomico di Brera, Milan, Italy)
- K. Makishima (Department of Physics, University of Tokyo, Tokyo, Japan)
- L. Maraschi (INAF-Osservatorio Astronomico di Brera, Milan, Italy)
- K. Pounds (Leicester University, Leicester, UK)
- M. Rees (Institute of Astronomy, Cambridge, UK)
- G. Setti (INAF Istituto di Radioastronomia, Bologna, Italy)
- R. Sunyaev (Max-Planck-Institut fuer Astrophysik, Garching, Germany)
- Y. Tanaka (Max-Planck-Institut for Extraterrestrial Physics, Garching, Germany)
- H. Tananbaum (Harvard Smithsonian Center for Astrophysics, Cambridge, USA)
- J. Trümper (Max Planck Institute for Extraterrestrial Physics, Garching, Germany)
- C. M. Urry (Yale University, Yale, USA)

Local Organizing Committee

- V. Braito (INAF-Osservatorio Astronomico di Brera, Milan, Italy)
- P. Caraveo (INAF-IASF, Milan, Italy)
- R. Della Ceca (INAF-Osservatorio Astronomico di Brera, Milan, Italy)

A. De Luca (INAF-IASF, Milan, Italy)

- P. Esposito (INAF-IASF, Milan, Italy)
- F. Gastaldello (INAF-IASF, Milan, Italy)
- R. Millul (INAF-Osservatorio Astronomico di Brera, Milan, Italy)
- G. Pareschi (INAF-Osservatorio Astronomico di Brera, Milan, Italy)
- A. Tiengo (INAF-IASF, Milan, Italy)
- F. Tavecchio (INAF-Osservatorio Astronomico di Brera, Milan, Italy)
- G. Trinchieri (INAF-Osservatorio Astronomico di Brera, Milan, Italy)

464

Organizing Institutions:

Istituto Nazionale di Astrofisica (INAF) INAF – Osservatorio Astronomico di Brera INAF – IASF-Milano

Contributions from:

SAIt - Società Astronomica Italiana Thales Alenia Space Italia

Support from:

ASI - Agenzia Spaziale Italiana CIFS - Consorzio Interuniversitario per la Fisica Spaziale Museo Nazionale della Scienza e della Tecnologia Leonardo da Vinci

Thanks to:

F. Ribes Bertoneu, L. Reduzzi, L. Ronzon, F. Galli (National Museum of Science and Technology)

G. Malaspina, C. Zazzetta (INAF)

A. Wolter, L. Stella (unofficial photographers)

Special Thanks to: D. Fugazza and R. Panzera for the invaluable and continuous support (web pages, database management, macros, shortcuts and so on) at all times, from before the conference to the end of the volume editing.

FOREWORD

The discovery of the first X-ray source outside the solar system (Sco X-1) by Riccardo Giacconi and his group, in 1962, initiated a new era in astronomy. For the first time it was realized that celestial bodies could emit enormous powers outside the traditional optical wavelength range.

In its first 50 years, X-ray astronomy has provided a new vision of the universe, with unsuspected high-energy activities at all scales.

In our Galaxy, a wide variety of systems: rapidly rotating Neutron Stars - remnants of Supernova explosions - compact objects accreting matter in binary systems - a Super Massive Black Hole lurking at the Galaxy's centre - emit X-rays carrying fundamental information on their nature and evolution. Similar sources are found in external galaxies, and can be traced back in time, to witness their collapse and merging history.

On cosmological scales, the intergalactic gas in clusters of galaxies shines at X-ray energies allowing us to probe the gravitational field in these systems, related to the original seeds of structure in the early Universe. Super Massive Black Holes at the centers of galaxies, fed by the infall of gas, release extreme luminosities, that can be traced to very large distances, carrying information on high energy processes near the Black Hole horizon as well as on the evolution of the Universe.

In these five decades, astronomical instrumentation at other wavelengths, both from ground and from space, has also seen a fantastic progress in capabilities and in the opening of new observing windows, strongly driven by questions posed by the High Energy Universe and by Cosmology at large.

Understanding the physics of the High Energy Universe needs a fundamental step forward in space projects, reaching fainter fluxes and resolving structures across a large field of view with improved timing and spectroscopic capabilities over a wide energy range.

The Conference held in Milan in October 2012, on the occasion of the 50th birthday of X-ray astronomy, was conceived to bring together the fundamental knowledge and experience accumulated in the past, recognize the current scientific and technological achievements of X-ray astronomy today with a view to envisaging the scientific challenges for the next 50 years.

The meeting was well attended. Close to 200 scientists participated in an atmosphere of open discussions and collaboration. Some of the most representative scientists from around the world were among the participants. Lord Martin Rees, Astronomer Royal, gave an inspiring opening talk. The Nobel Prize Riccardo Giacconi, Joachim Trümper and Jasuo Tanaka Johan Bleeker, Mike Watson and Luigi Piro presented the pioneering missions that were developed in the early days of X-ray astronomy.

Weather and location, with a wide cloister hosting lunches and coffee breaks, collaborated in favoring exchanges of ideas between scientists of different countries and generations notwithstanding animated discussions. The enclosed pictures may give a hint of the meeting atmosphere.

With the present volume we hope to provide the community at large with a testimonial of the essence of the meeting. We thank all speakers, authors, participants and supporters for their contributions that helped to make the event pleasant and successful.

G. Trinchieri, R. Della Ceca, T. Maccacaro, L. Maraschi, G. Pareschi









